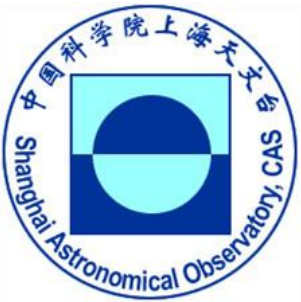




3D dust distribution in the Galaxy based on LAMOST

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Supervisor : Shiyin Shen

OUTLINE

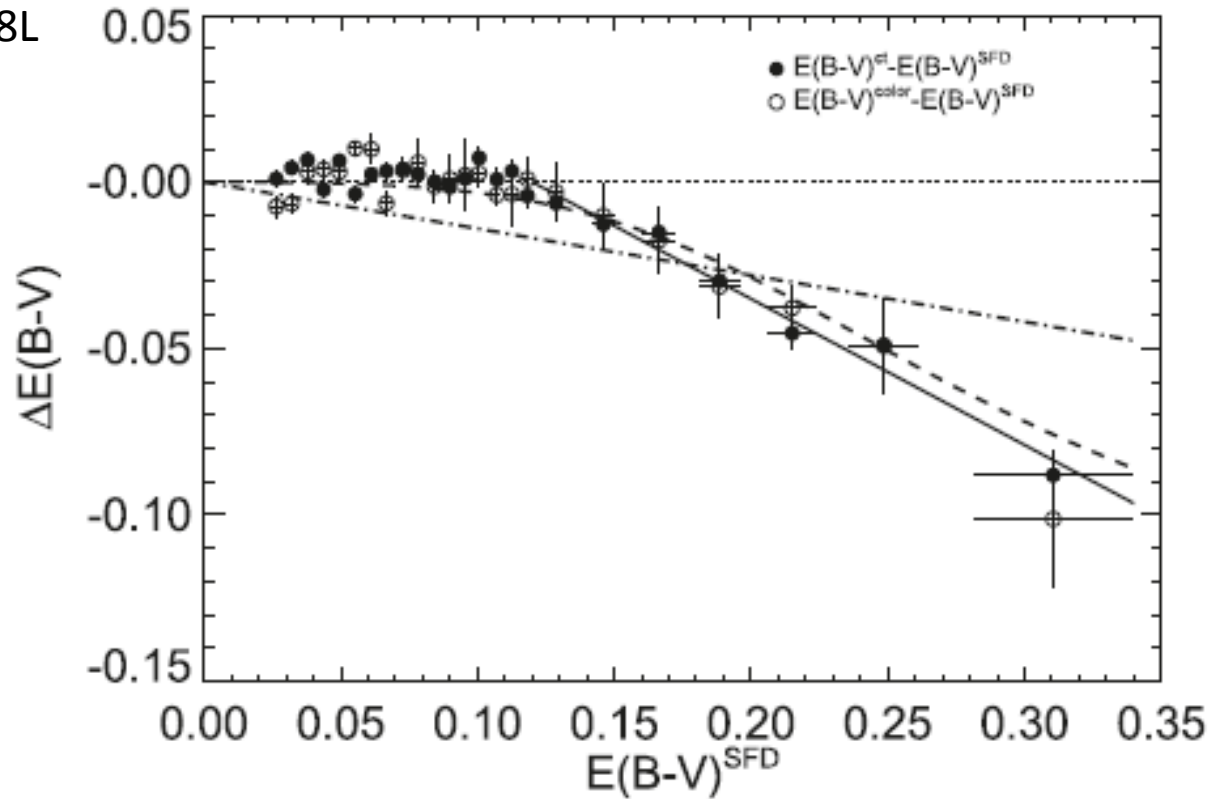
- **background**
- **Data and method**
- **Primary result**
- **Future work**

background

➤ 2D Galactic extinction based on SCUSS

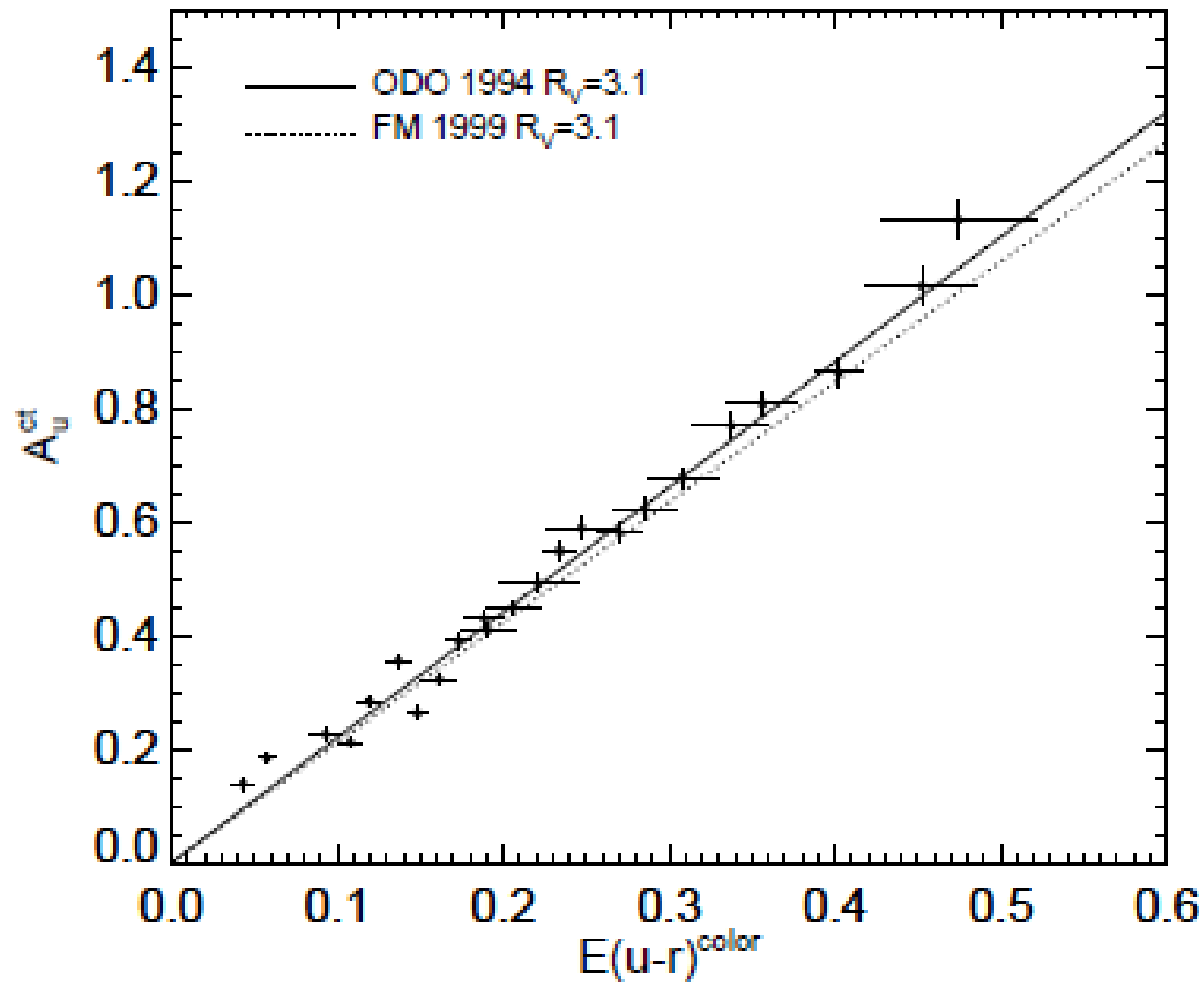
Li et al. 2017 AJ 153 88L

$$b < -30^\circ$$



$E(B - V)_{SFD} < 0.12 \text{ mag}$: the extinction and reddening from the number counts and color distribution are **in good agreement with** the SFD results.

$E(B - V)_{SFD} > 0.12 \text{ mag}$: the SFD map **overestimates** the Galactic reddening systematically, $\Delta E(B-V) = 0.43[E(B-V)_{SFD} - 0.12]$.



The Galactic extinction curve is in good agreement with the $R_V = 3.1$ extinction law of **O'Donnell (1994)**.

➤ Many works on 3D extinction map.

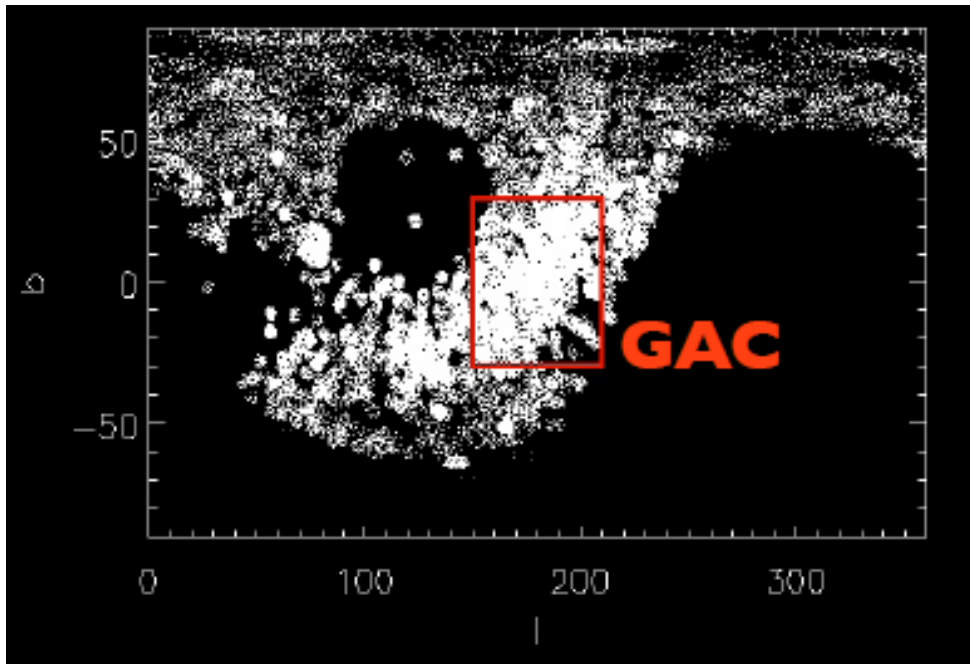
(Sale et al.2014; Chen et al. 2014; Hanson et al. 2016)

➤ What's the smooth structure of dust distribution?

	scale length (kpc)	scale height (pc)
Drimmel et al.2001	2.26	134
Sharma et al. 2011	4.20	88
Pandey et al. 1987	---	160
Joshi 2005	---	125

● Data

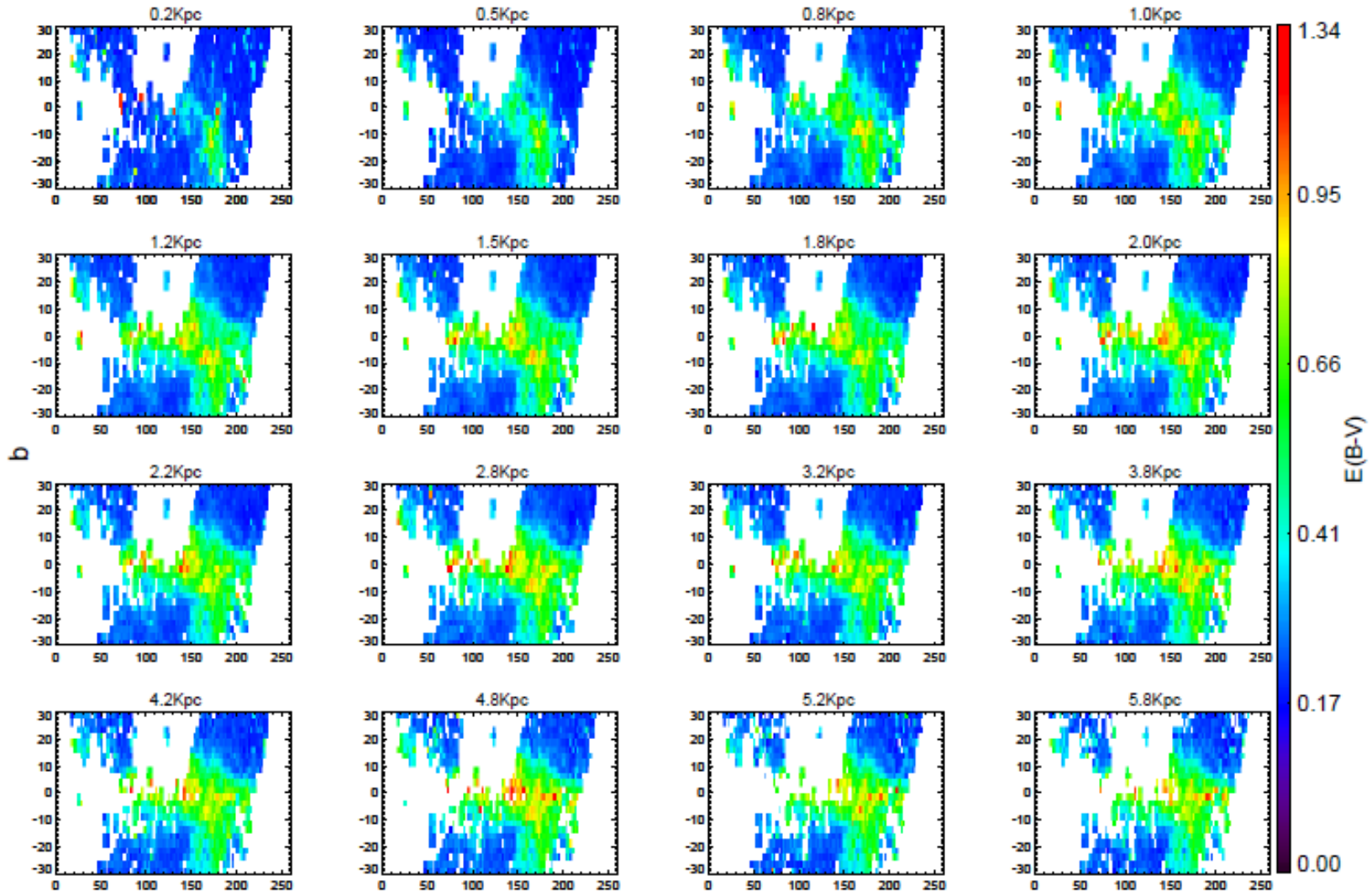
6M stars with independent, accurate measurement of $E(B-V)$ and distance by Yuan et. al.



$$-30^\circ < b < 30^\circ$$

● method

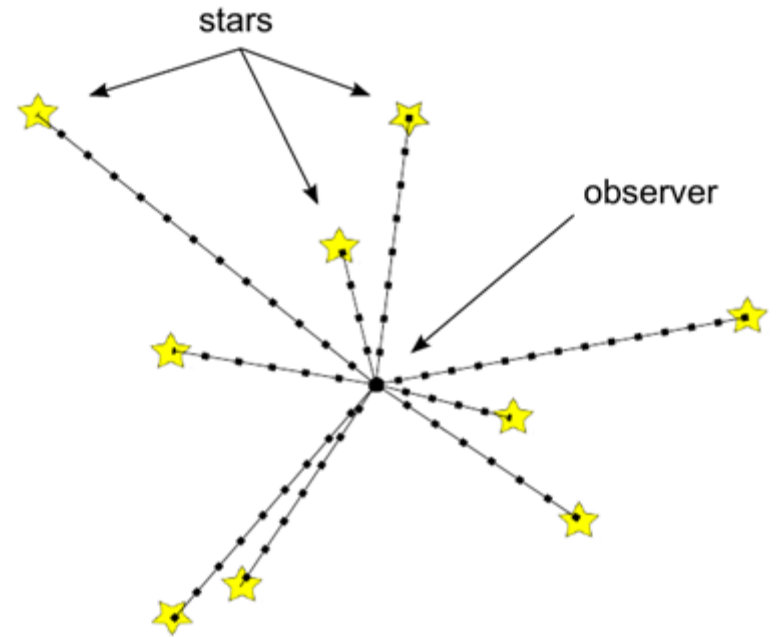
$$E_i^{observe}(l_i, b_i, r_i) : 2.5^\circ \times 2.5^\circ \times 250 pc$$



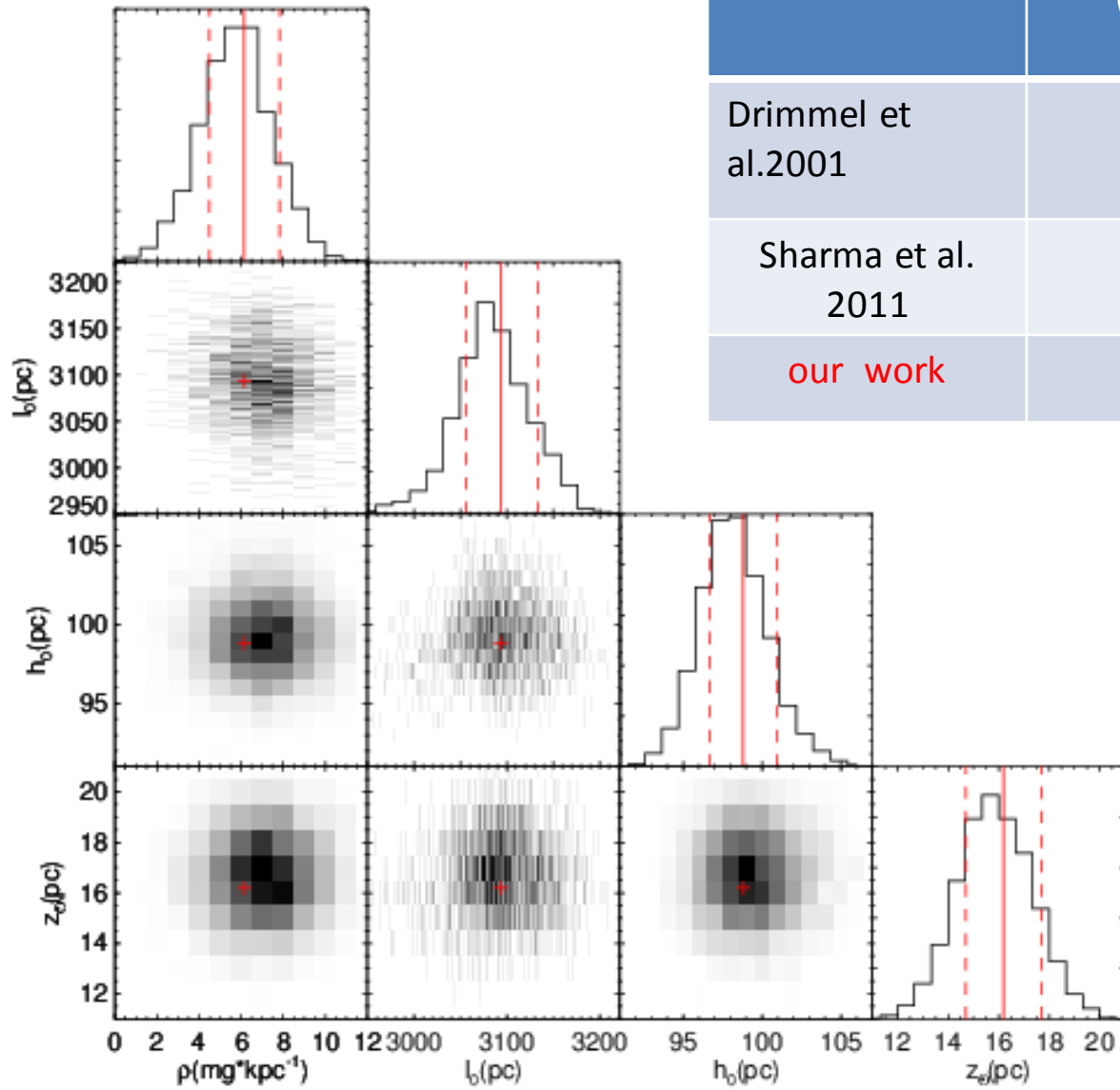
$$E_i^{\text{model}} = \int_0^\infty \rho_{\text{Dust}}(l_i, b_i, r_i) dr$$

$$\rho_{\text{Dust}}(l, b, r) = \rho_0 \exp\left[-\frac{R}{l_0} - \frac{|z|}{h_0}\right]$$

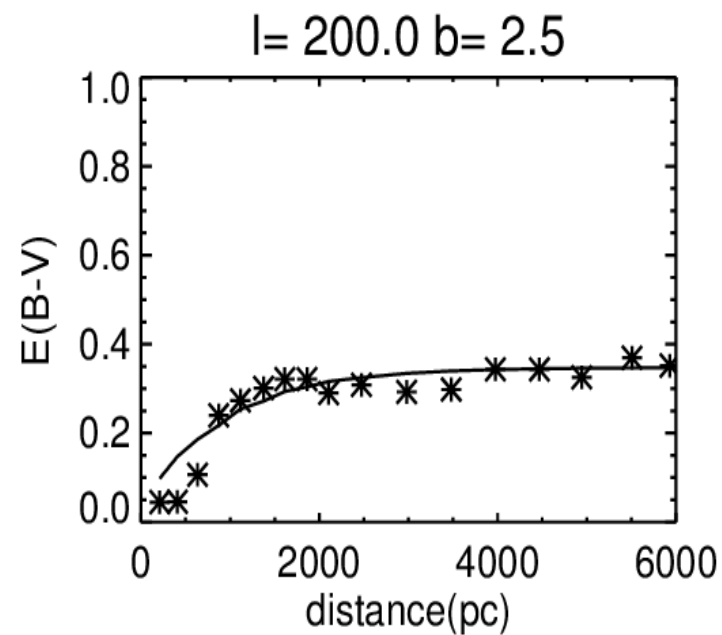
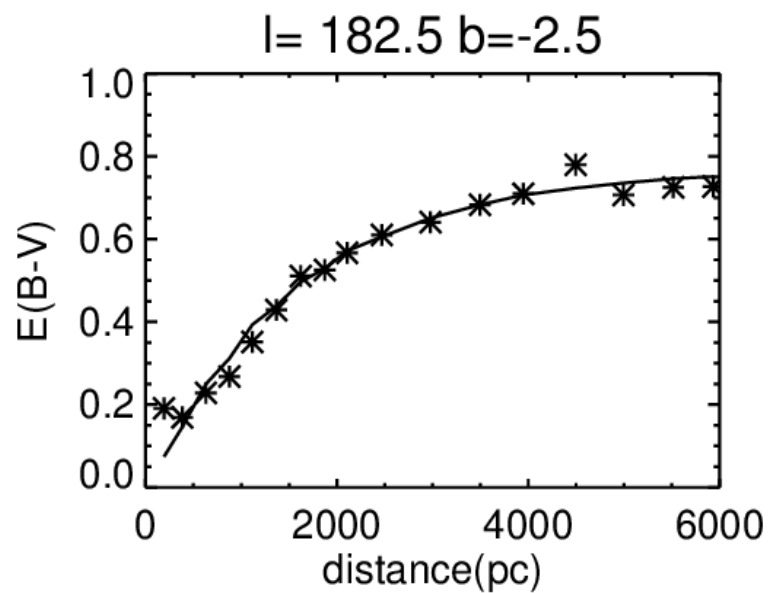
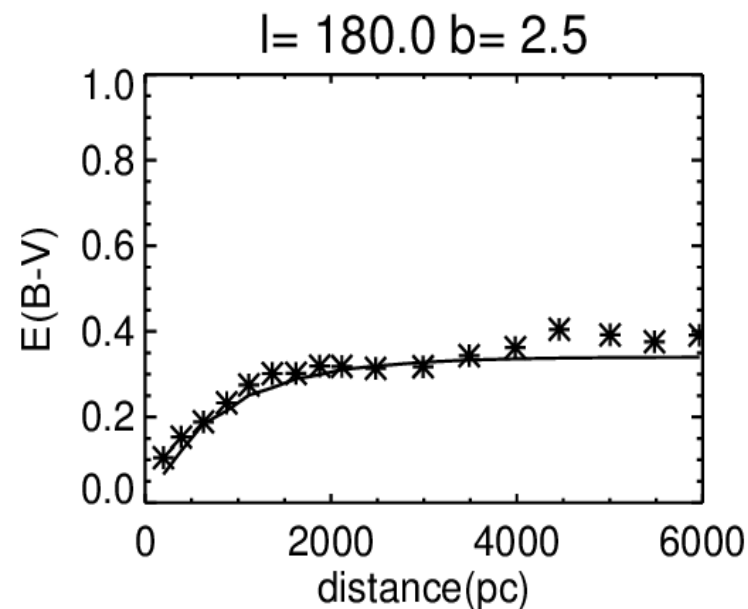
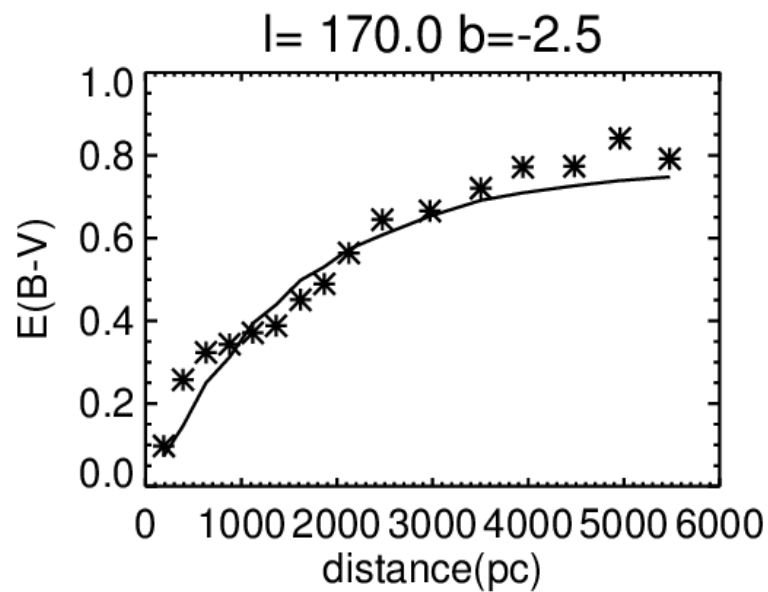
MCMC fit : $\rho_0, l_0, h_0, Z_\odot$



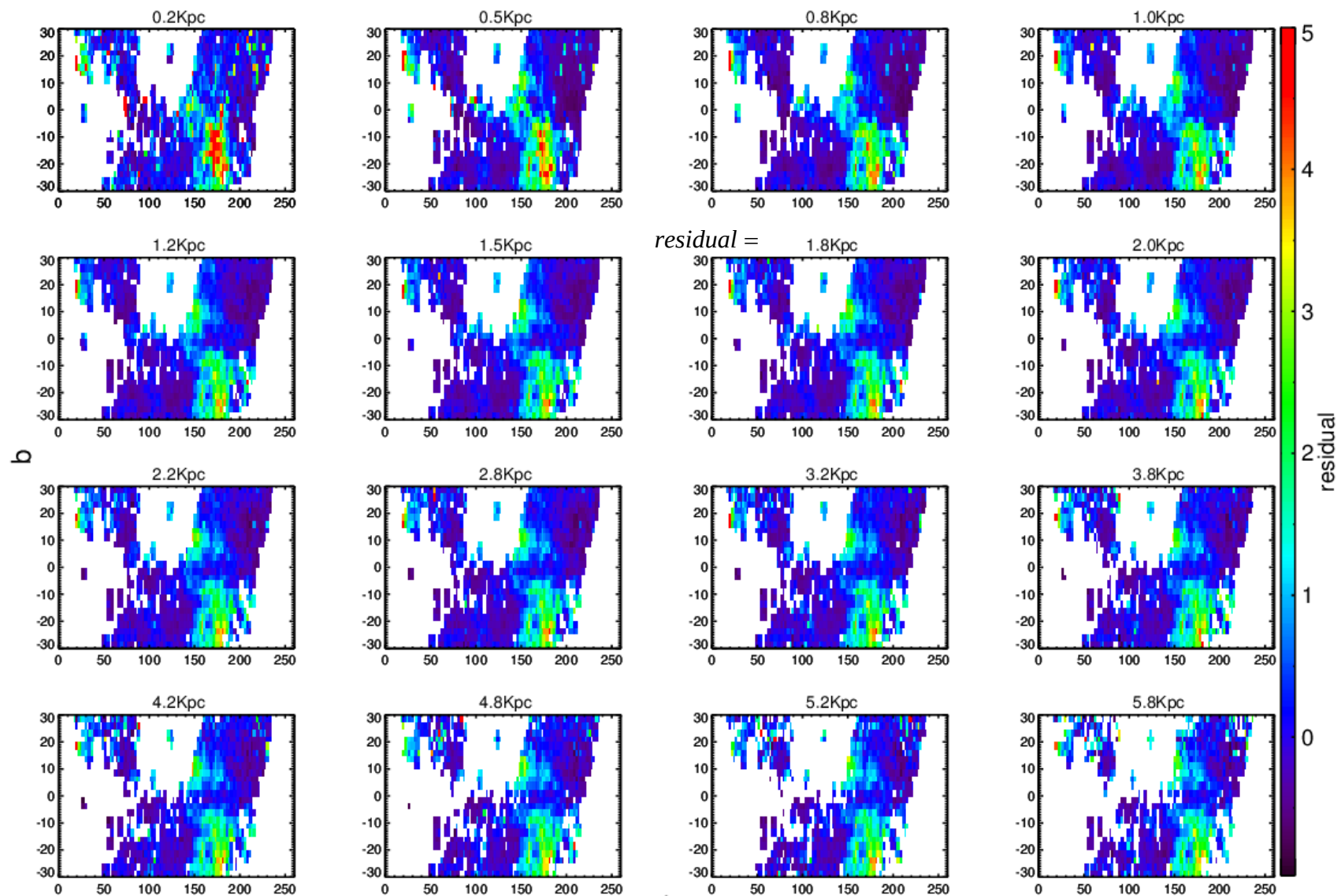
Primary result

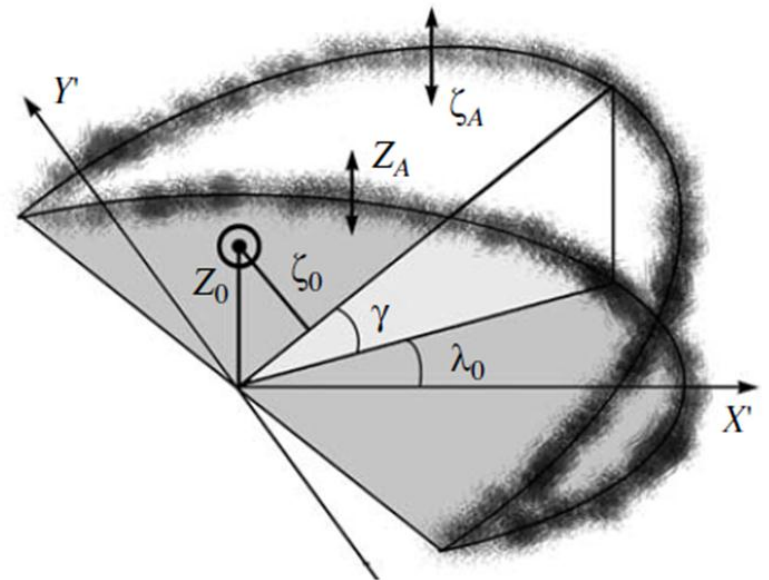
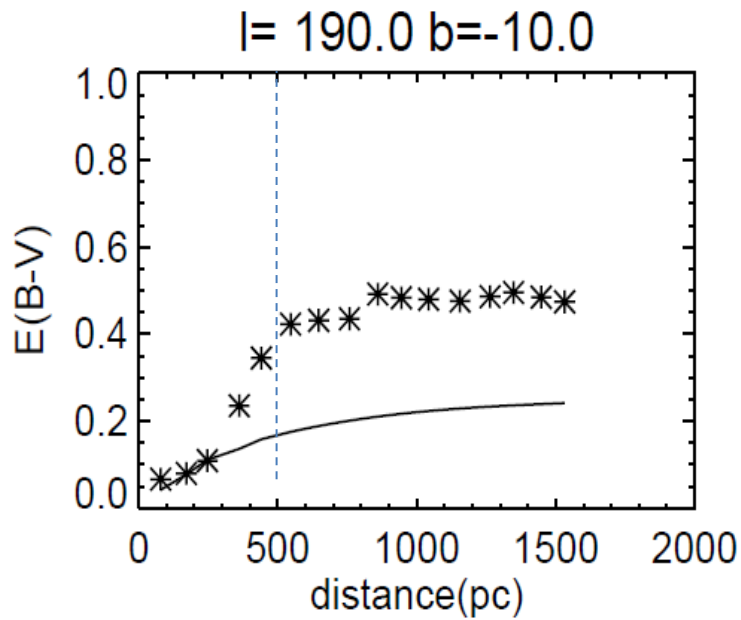
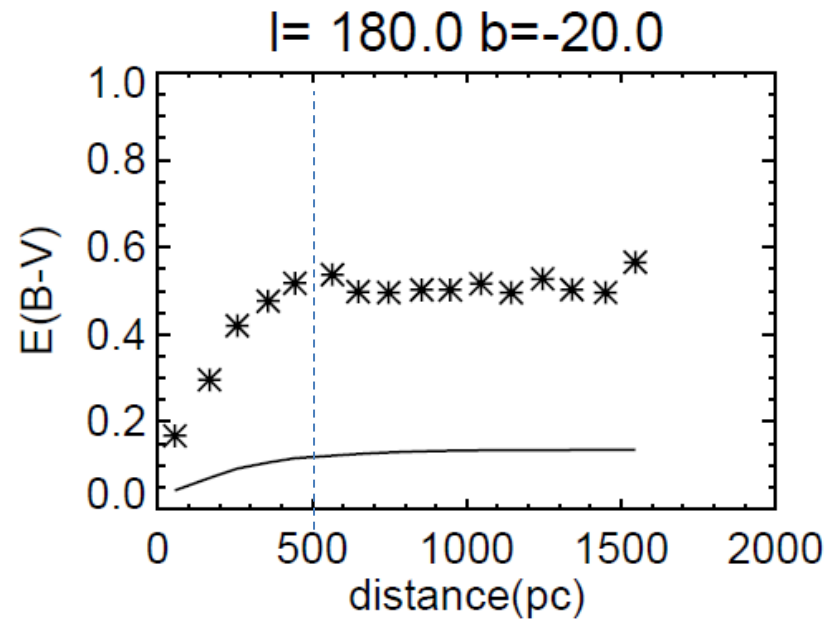
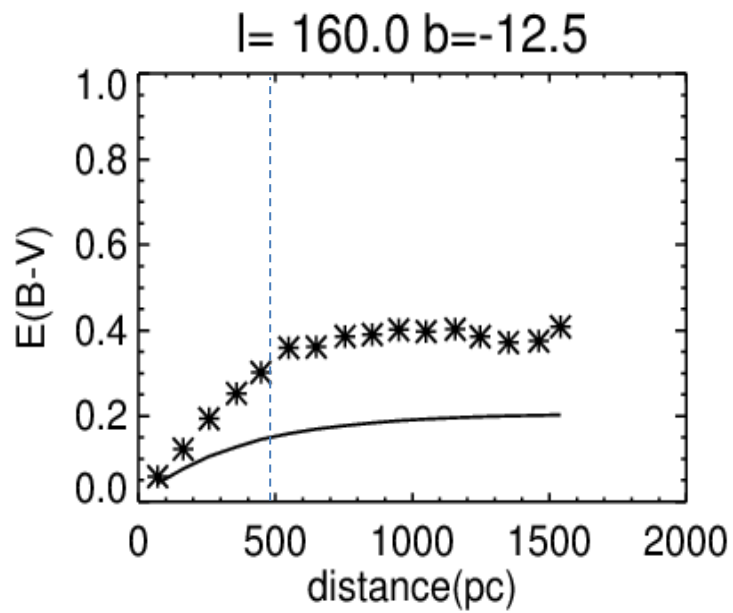


	Scale length (kpc)	scale height (pc)
Drimmel et al.2001	2.26	134
Sharma et al. 2011	4.20	88
our work	3.09	98



$$residual = (E_{ob} - E_{mo}) / E_{mo}$$

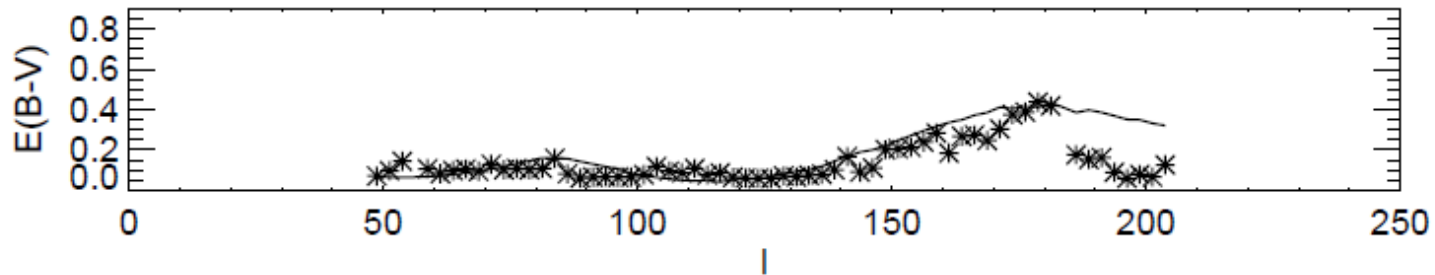




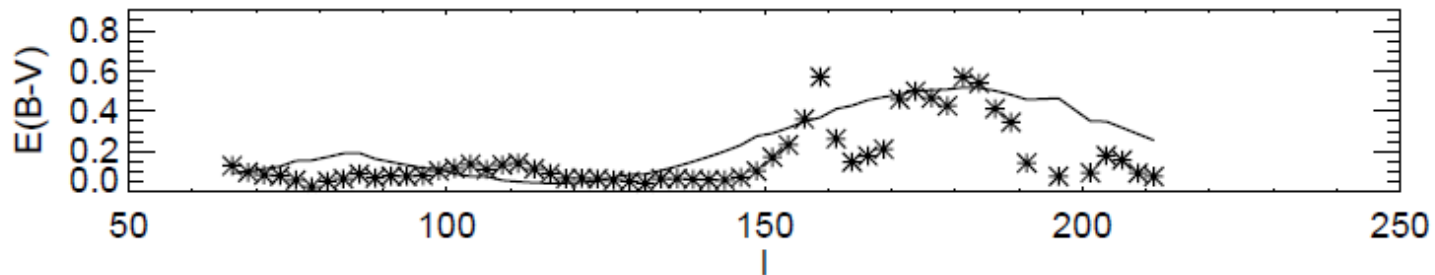
● Future work

500pc

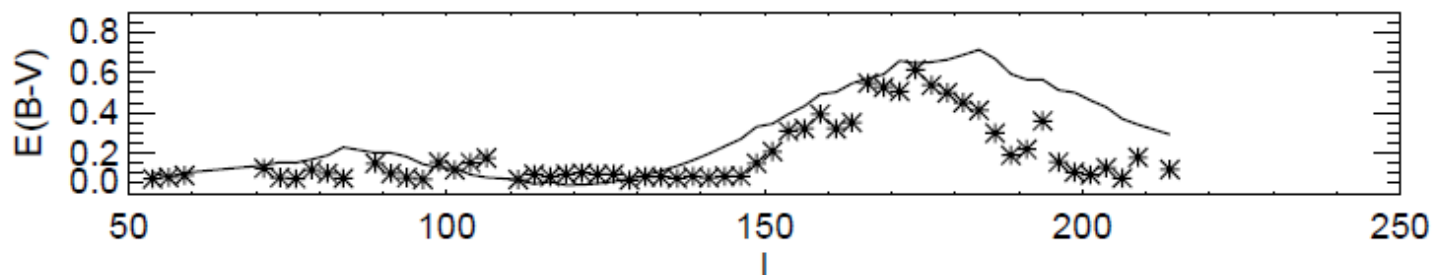
$b = -25$



$b = -20$



$b = -15$



Thank you !